



Gravitino production from WW scattering in the broken phase

$$W^a + W^b \longrightarrow \widetilde{W}^c + \widetilde{G}$$

$$W^+ + W^- \longrightarrow \widetilde{\chi}_i^0 + \widetilde{G}$$

(in other words, "About the mass as a source of Evil") <u>Andrea Ferrantelli</u>

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<0|Motivation: why do we study the gravitino?|0>

- Cosmological interest in the gravitino has grown because ... -
- The gravitino \widetilde{G} is the gauge field of local SUSY transformations, therefore of SUGRA theories. It is the supersymmetric partner of the graviton, with mass provided by the super-Higgs mechanism.
- In a number of scenarios (e.g. in gauge mediation), the gravitino is the <u>lightest supersymmetric particle</u> (LSP). If R-parity is conserved, we then have the immediate consequence that such a particle is a very attractive candidate for Dark Matter.
- If it is not the LSP (e.g. in gravity mediation), it may decay during the Big Bang Nucleosynthesis (BBN), causing EM or hadronic showers, which <u>disintegrate the primordial nuclei</u> (this is one of the aspects of the "gravitino problem").

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<0 WW scattering: the basis of gauge eigenstates |0> Within the particle content of the MSSM, one can consider gluon scattering into gluino and gravitino:

$$g^a(k) + g^b(k') \longrightarrow \tilde{g}^c(p') + \tilde{G}(p)$$

which contributes to the thermal production of gravitinos in the early universe (M. <u>Bolz</u>, A. <u>Brandenburg</u>, W. <u>Buchmueller</u>, Nucl. Phys. B 606 (2001) 518).



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Let us now take into analysis the following Electroweak process in supergravity, with a wino and a gravitino in the final state:

$$W^a + W^b \longrightarrow \widetilde{W}^c + \widetilde{G}$$

The corresponding Feyman diagrams are analogous to those in the gluon scattering, since the couplings in supergravity are <u>universal</u>. Thus the amplitude is formally the same in both WW and gluon scattering. The only difference is found in the kinematics and in the <u>spin sums</u>.



FIG. 2: The four diagrams which contribute to $W^a + W^b \longrightarrow \widetilde{W}^c + \widetilde{G}$

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• <u>High energy limit</u> (at energies which are well above the scale of Electroweak breaking, the W can be considered <u>massless</u>):

$$\sum_{spin} |\overline{\mathcal{M}}|^2 = \frac{1}{9} \sum_{spin} |\mathcal{M}|^2 = 4 \frac{g^2 |\epsilon^{abc}|^2}{9M_P^2} \left(1 + \frac{m_{\widetilde{W}}^2}{3m_{\widetilde{G}}^2}\right) \left(s + 2t + 2\frac{t^2}{s}\right)$$

(J. Pradler and F. D. Steffen, Phys. Rev. D 75 (2007) 023509)

• Low energy limit (below the EW scale of 100 GeV, W is massive):

$$\sum_{spin} |\overline{\mathcal{M}}|^2 = 4 \frac{g^2 |\epsilon^{abc}|^2}{9M_P^2} \left[\left(1 + \frac{m_{\widetilde{W}}^2}{3m_{\widetilde{G}}^2} \right) \left(s + 2t + 2\frac{t^2}{s} \right) - \frac{t(s+t)}{3m_{\widetilde{G}}^2} + \mathcal{O}\left(\frac{m_i^2}{m_{\widetilde{G}}^2} \right) f(s,t) \right]$$

Anomalous contribution $\left(\frac{d\sigma}{dt}\right)_{cm} \approx \frac{g^2 |\epsilon^{abc}|^2}{64\pi M_P^2} \frac{(1-\cos^2\theta)}{3m_{\widetilde{c}}^2}$

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 $\sigma_{tot} \approx$

<0| WW scattering: the basis of gauge eigenstates |0> <u>What happened</u>?

By directly using the spin sums (1) and (2), we can separate the contributions of the <u>Transverse and Longitudinal degrees of freedom of the gauge bosons</u>, both in the high energy limit and in the broken phase:



Andrea Ferrantelli – Gravitino production from WW scattering in the broken phase <0 WW scattering: the basis of gauge eigenstates |0> What is missing? Should we add the contribution of the Higgs? The rise of a problem: the coupling terms between gravitino and supercurrent in the general Supergravity Lagrangian

$$\mathcal{L}_{\psi J} = -\frac{i}{\sqrt{2}M_{Pl}} \left(\widetilde{\mathcal{D}}_{\nu} \phi^{*j} \bar{\psi}_{\nu} \gamma^{\mu} \gamma^{\nu} \chi_{L}^{j} - \widetilde{D}_{\nu} \phi^{j} \bar{\chi}_{R}^{j} \gamma^{\nu} \gamma^{\mu} \psi_{\nu} \right) \\ -\frac{i}{8M_{Pl}} \bar{\psi}_{\mu} \left[\gamma^{\nu}, \gamma^{\rho} \right] \gamma^{\mu} \lambda^{(a)} F_{\nu\rho}^{(a)},$$

do not admit any scalar-gravitino-gaugino vertex.

We must then use the method of mass insertions:



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mixing of winos and higgsinos in the t and u channels: $M^{(1)}$

diagrams with two mass insertions in the external leg:





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Consequently, the total amplitude has the following expression:

$$\mathcal{M}(W^a + W^b \longrightarrow \widetilde{W}^c + \widetilde{G}) = M^{(0)} + M^{(1)} + M^{(H)} + M^{(2)}$$

where the original amplitude, with no mass insertions, has been labeled with $M^{(0)}$



The Higgses contribute to the original anomaly and provide with new divergences. It has been shown that there is no cancellation of the anomalies found in the scattering of massive W bosons, neither by including the Higgs and considering higher orders in a perturbation theory developed around a mass parameter. Within the MSSM, there is no other suitable particle.

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Conclusions and perspectives

We have calculated WW scattering in Supergravity (SUGRA), namely:

$$W^a + W^b \longrightarrow \widetilde{W}^c + \widetilde{G}$$
 $W^+ + W^- \longrightarrow \widetilde{\chi}_i^0 + \widetilde{G}$

• It turns out that, in the squared amplitude, unitarity is <u>preserved at any scale</u> only if the W bosons are considered <u>massless</u>. In they are <u>massive</u>, new terms, the so-called "anomalies" which would <u>violate unitarity</u>, are generated.

The result holds in both the bases of gauge and mass eigenstates, due to the <u>non</u> vanishing mass of the W boson, which affects the longitudinal modes of the Ws and their couplings with the gravitino. Moreover, <u>in the MSSM the Higgs bosons</u> <u>do not cure the problem</u>, but contribute with new divergences.

• We may need to understand this result more deeply, first of all from a formal point of view, as it is <u>unexpected</u> in the context of SUGRA theories and may provide with interesting theoretical issues.

• From a more phenomenological perspective, this production channel of gravitinos can be observed as a secondary reaction at the <u>LHC</u> (through e.g. gluon-fusion), with the related implications in the context of <u>Cosmology</u>.

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- Work in Progress!



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